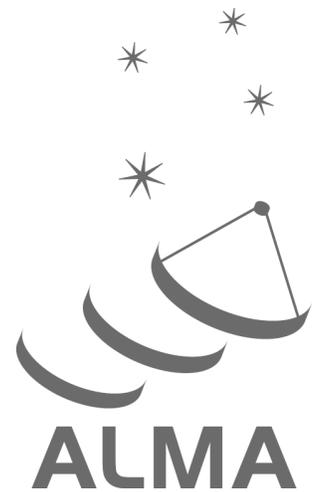


Guide to the European ALMA Regional Centre



www.almascience.org

User Support:

For further information or to comment on this document, please contact your regional Helpdesk through the ALMA Science Portal at www.almascience.org. Helpdesk tickets will be directed to the appropriate ALMA Regional Centre at ESO, NAOJ or NRAO.

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Contributors

Martin Zwaan, Evanthia Hatziminaoglou, the European ARC at ESO and the European ARC nodes.



EUROPEAN ARC
ALMA Regional Centre

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Purpose of this document

The purpose of this document is to help users to find their way around the European ALMA Regional Centre (ARC). The European ARC is the 'one stop shop' for all European ALMA users. This document explains how the ARC structure works and how it can help users at the various stages of their ALMA observing project.

ALMA in a nutshell

The Atacama Large Millimetre/submillimetre Array (ALMA) is expected to be the leading observatory at millimetre and submillimetre wavelengths for many decades to come. ALMA is a global collaboration involving East Asia, Europe, North America and the host country Chile. ALMA comprises 66 high-precision antennas equipped with receiver and digital electronics systems to observe in the frequency range from 30 GHz to 1 THz and at angular resolutions down to 6 milliarcsec. Using a fully dynamic scheduling system and innovative calibration strategies, the ALMA system allows observers to make the best use of the atmospheric conditions on the Chajnantor plateau at 5000 m altitude in the Atacama desert. The European contribution to the ALMA project is led by ESO on behalf of its member states.

ALMA has performed scientific observations since September 2011, in addition to regular commissioning and science verification. Each observing period offers another opportunity for astronomers to propose science projects with ALMA. In the call for proposals for Cycle 5 the following capabilities are offered: forty three 12-m antennas in the 12-m Array and eleven 7-m antennas in the Morita array, also called Atacama Compact Array (ACA) plus four 12-m antennas for single-dish observations (Total Power Array, which can also be added to the main array); eight frequency bands (3, 4, 5, 6, 7, 8, 9, and 10, covering wavelengths from about 3.1 to 0.35 mm) and maximum baselines ranging from 155 m to 16 km; band 5 observations from March 2018 on, band 4 polarisation observations, VLBI observations, Solar Observations and Large Programmes. More details on the capabilities of ALMA are given in the Proposer's Guide, available at the ALMA Science Portal at <http://almascience.org>. More information on the ALMA project can also be found on the Science Portal.

The ALMA Regional Centres (ARCs)

The interface between ALMA and the astronomical community is provided by the three partners through the ALMA Regional Centres (ARCs). These are located at NAOJ in Mitaka, Japan, for the East Asian partnership, at ESO in Garching, Germany, for the European partnership, and at NRAO in Charlottesville, USA, for the North American partnership. The ARCs are staffed by scientists with expertise in radio astronomy, millimetre and submillimetre astronomy, and interferometry, and their purpose is to work with the community of astronomers to maximise the scientific productivity of the telescope.

Organisation of the European ARC structure

The ARC structure in Europe supports the scientific use of ALMA by the European user community. The ARC forms the interface between the ALMA observatory and the user community throughout the lifetime of a project, i.e., from proposal preparation to data analysis. In Europe, the services to the community are provided by a distributed network of ARC nodes at Bologna, Bonn/Cologne, Grenoble, Leiden, Manchester, Ondrejov, and Onsala, as well as an ALMA Centre of Expertise (CoE) in Lisbon, under the coordination of the ESO ARC in Garching.

The primary aim of the European ARC is to maximise ALMA's scientific return for European users. Each ARC node is staffed with scientists providing a range of experience in interferometry, (sub-)mm observing and ALMA data reduction and interpretation, as well as a range in scientific interests. The staff at the ESO ARC, the ARC nodes and the CoE work together to provide optimal support to users during the complete lifetime of a project from proposal preparation, choice of observing modes, preparation of the Phase 2 material, delivery of the calibrated science products to the users, and, if required, additional data reduction support. Furthermore, the ARC can give advice regarding observing strategies and can help with ALMA archival research. The ARC network organises community days, tutorials and workshops throughout Europe to reach out to the community and to train users in proposal preparation and data reduction. Several ARC nodes also supervise Master's and PhD students and give university lectures on ALMA-related topics.

The central point of contact for current and prospective ALMA users is the ALMA Helpdesk, which is available to all registered users from the *ALMA Science Portal*. There is one Helpdesk for all ALMA users, but queries from European users will be primarily addressed by the European ARC, either by scientists in Garching or by staff in one of the ARC nodes. The Helpdesk includes a library of “Knowledge-base” articles that address a number of common issues and questions. If the answer is not found in the suggested Knowledge-base articles, the user can submit a ticket. During normal operations, users can expect a response within two business days. During the week prior to a proposal deadline, extra staff will be assigned to the Helpdesk and every effort will be made to address time-critical questions in a rapid manner.

For all ALMA questions: use the ALMA Helpdesk to contact the ARC. The Helpdesk is available through the [ALMA Science Portal](#).

Role of the central ARC, the ARC nodes and the CoE

The central ARC in Garching and the ARC nodes distributed throughout Europe work together to provide optimal support to European ALMA users. It is however important to realise that the roles that the central ARC, the nodes and the CoE fulfil are very different. For example, all face-to-face support, one of the core functions of the ALMA project, takes place at the ARC nodes and the CoE.

The different roles of the ARC and the ARC nodes are outlined below. Note that the typical ALMA user does not need to know the specifics of how tasks are divided between the central ARC and the nodes: All initial contacts between the user and the ARC staff happen through the ALMA Helpdesk. The ARC staff may answer the user's question immediately, refer to documentation, or forward the query to an expert within the ARC-nodes network. The ARC may advise that face-to-face help is required, or the users may indicate themselves that they need to get in touch with their Contact Scientist (CS) or request face-to-face support.

The main tasks of the central ARC, of direct relevance to ALMA users, are:

- Phase 1 operations: distribution of the call for proposals to prospective European ALMA users, user support during proposal preparation as well as assistance in coordinating the refereeing process.
- Phase 2 operations: inspection of the Scheduling Blocks and implementation of changes following successful change requests.
- Data product support: quality assurance (QA2) of science data sets and delivery of the final raw and reduced data to the PIs.
- Archive operations: the ESO ARC node holds a complete synchronised copy of the ALMA archive.
- Running the ALMA Helpdesk.

The relevant tasks of the ARC nodes are:

- Quality Assurance (QA2) of science data sets.
- User support with proposal preparation by means of face-to-face meetings, dedicated workshops and through the ALMA Helpdesk.
- Phase 2 operations: assisting users with the preparation of their Phase 2 material, and the technical details required to schedule and execute the proposed observations, as well as with the preparation and validation of their Phase 2 material.
- Face-to-face help with data reduction, including expert support in data processing for specialised observing techniques.
- Help in archival research, including assistance to users of the ALMA archive in identifying and using the data products suitable for their scientific projects.
- Facilitate science coordination between ALMA users, i.e. providing space for small meetings, putting people with similar interests in contact if mutually desired etc.
- Close interaction with the community and regular updates on ALMA (Community Days and other events, newsletter, webpages), as well as community preparation through lectures, tutorials, etc.

- Outreach to the general public.

The tasks of the ALMA Centres of Expertise:

European ALMA Centres of Expertise are a recent addition to the European ALMA support structure. A CoE has a temporary status and may eventually acquire the status of full ARC node, once the conditions described in detail in the document “European ALMA Centres of Expertise: Tasks and Responsibilities” have been satisfied. The tasks of the CoEs are similar to those of the ARC nodes, with the exception that they do not provide CSs or face-to-face support for data reduction. Once a CoE becomes an ARC node, its role and responsibilities will fully align with those of the other ARC nodes.

The following section describes in more detail the kind of support to be expected throughout the lifetime of an ALMA project.

How the ARC can help with your ALMA observing project

For each scheduling period, a Call for Proposals for ALMA is issued. The ESO ARC distributes this call among the European user community by email and makes it available on the ALMA Science Portal (<http://almascience.org>). The call provides information about the available capabilities for each cycle, necessary information for the electronic submission of proposals as well as links to documentation such as the Proposer's Guide, the ALMA Technical Handbook and the Observing Tool (OT) User and Reference Manuals, as well as the OT Quick Start Guide. For the complete list of documentation, see the relevant section on the ALMA Science Portal.

The ALMA Science Portal

All ALMA users need to register to the ALMA Science Portal. This is a single sign-on gateway to the various tools described below, the Helpdesk, the ALMA Science Archive, and other relevant information. Each astronomer who uses ALMA is assigned a single ARC for user support at the time they register. ALMA users should have one ALMA account, that they should update in case any of their profile details changes (e.g. e-mail address, affiliation etc).

To submit a proposal, download ALMA software, or gain access to the Helpdesk, you will need to register with the [ALMA Science Portal \(almascience.eso.org/\)](http://almascience.eso.org/).

The proposal stage

The Phase 1 material of the proposals must be prepared with the ALMA OT which is provided by the Joint ALMA Office (JAO) in Chile and is available from the ALMA Science Portal, section "Tools". The OT requires the users to provide a scientific and technical justification, target specification, and sensitivity. The OT provides the users with time-on-target plus overhead specification, integration time estimation, transparency and atmospheric conditions, etc.

Use the [ALMA Observing Tool](#) to prepare your proposal. Need help? Use the [Helpdesk](#) to contact the ARC.

The ALMA Observing Support Tool (OST), hosted at the UK node (<http://almaost.jb.man.ac.uk/>) provides an easy web-based interface to preparing simulations. The user can upload a FITS image and 'observe' it with ALMA at different frequencies, bandwidths, configurations, integration times and weather conditions. The results are images as ALMA would

produce them, together with point spread functions and maps of the *uv*-coverage. This information can help users in planning their ALMA observations.

More experienced users may want to use simulation capabilities included in the CASA data reduction package, which have more flexibility. For more details see <http://casaguides.nrao.edu>.

Staff at the EU ARC network assist users with the preparation of their Phase 1 material. This assistance covers help with the Observing Tool, the ALMA exposure time calculator and the simulator. Staff are also available for consultation regarding observing strategies.

Users may also choose to visit an ARC node for face-to-face support during proposal preparation.

Need help with proposal submission? Use the [Helpdesk](#). Need face-to-face help for preparing a complicated proposal? Use the [Helpdesk](#) to coordinate a visit to an ARC node.

After the submission deadline, all proposals are ranked by the ALMA Proposal Review Committee (APRC), which makes a recommendation to the ALMA director on the scheduling of ALMA projects for the coming period. Note that there is only one international proposal review committee for all ALMA proposals.

The preparation of the Phase 2 material

In Cycle 5, PIs will be generating their Phase 2 Science Goals (SGs). PIs will be notified by email on the ranking of their project as well as on the timeline for the submission of their Phase 2 material. The procedure for PIs to create their Phase 2 SGs is described in detail in the Proposer's Guide, available via the ALMA Science Portal. PIs can request assistance with the verification of their SGs from their Contact Scientist (CS), that will be assigned to all accepted A-, B-, and C-rated projects. Minor mistakes on the positions, velocities or frequencies can be corrected by the PIs directly. Major changes that might change the scientific content of the program or clash with another approved project, need special permission and should be requested via a special Helpdesk ticket.

Immediately after the Phase 2 deadline, the Scheduling Blocks (SBs) will be automatically generated by the observatory and will be released to ALMA operations in Chile for scheduling and eventual execution. No further alterations can be made without special approval, obtained via Helpdesk in consultation with a CS. Note that observers do not travel to Chile or elsewhere to take part in the observations - all ALMA observations are scheduled dynamically and executed in service mode.

The SBs can be executed at any time without further intervention from users, CS or ARC staff. If a problem is found at run-time, execution of the SB is halted. The Phase 2 team at the ESO ARC will fix the problem and the affected SBs are put back in the queue for scheduling.

Need the help of a Contact Scientist for preparing your Phase 2 material? Contact the ARC through the [Helpdesk](#).

Please note that in general PIs should submit their Phase 2 SGs as rapidly as possible to avoid any delays. Projects for which the Phase 2 material is not successfully submitted by the deadline given in the time allocation email will be downgraded. The Phase 2 team will make sure all SBs will run smoothly at the telescope (e.g. check calibrator variability, target proper motion).

Tracking your project

The Snooping Project Interface (SnooPI) allows PIs to follow the status of their ALMA observing projects after the project is accepted and scheduled for observing. SnooPI allows the PIs to search for their active projects and get an overview of the SBs that have been observed, those in the queue, those that are under processing, etc.

Follow project progress with [SnooPI](#), which is available on the Science Portal under the "Observing" menu.

After data have been taken

ALMA observing projects are dynamically scheduled, taking into account weather conditions, the configuration of the array, the proposal ranking and possibly other constraints. Short projects may be observed within one day, whereas others may take several weeks or months to complete. After the observations for a project have been completed, users are notified by email. Note that B- and C-rated projects will not be carried over to the next cycle when not completed, while A-rated projects will automatically be carried over.

All data pass through a multi-tier quality assurance (QA) scheme before they are released to the user. ALMA data sets can vary enormously in size depending on correlator settings, averaging times and total integration time. Data size may range from many TB for lengthy projects to tens of GB for the shortest observations. Users can download the data from the ALMA archive.

It is important to realise that *all* ALMA data are kept in the archive. All requests for data of any observing program are handled through this archive. This **ALMA Science Archive** is accessible through the Science Portal.

PIs can request, via the Helpdesk, the raw data of partly completed observations if further observations are expected to be taken in more than 90 days due to configuration constraints. Details on such intermediate data releases are given in the ALMA Users' Policies.

The data reduction pipeline

When a block of data for a particular program has been observed, the ALMA data reduction pipeline is activated. This pipeline runs on dedicated machines in Chile and in the ARCS uses a heuristics system to automatically edit, calibrate, and image data taken with the standard observing modes. The pipeline heuristics have captured the knowledge of experienced radio interferometer users and incorporated it into the reduction scheme. The output of the data reduction pipeline is stored in the ALMA Science Archive and is subject to a final Quality Assurance test (QA2) before the reduced data are made available to the user.

Pipeline calibration has been fully commissioned for most observing modes employed in previous cycles but about 25% of observing projects will still be calibrated manually using standard scripts (e.g., high frequency bands, full polarisation). The data processing will be done using a combination of the pipeline and manual analysis using CASA.

Data reduction support

The data reduction pipeline produces high quality science products for the standard observing modes. Expert hands-on help with data reduction may be required however, especially when more complicated observing techniques are used. The ARC is the primary contact point for users who need assistance with data reduction. ARC staff provide basic data analysis support, ranging from simple advice, to providing appropriate data analysis documents and products, to detailed assistance for users who require it.

The ESO ARC's primary responsibility is to ensure that the reduced data from all standard ALMA observing modes meet the requirements set by the approved observing proposal. The ARC nodes can provide further assistance to users who require it. A likely scenario may be the following: the ALMA data reduction pipeline has processed a user's data. If the data meet the requirements defined in the proposal, the data are delivered to the

*Basic data reduction support: use the [Helpdesk](#).
Advanced data reduction support: visit an ARC node (use the [Helpdesk](#) to coordinate a visit).*

user. However, the user may wish to make images with higher spatial resolution or higher sensitivity. The ARC nodes can help to reprocess the data, using different weighting schemes in the imaging stage. Another user may need higher dynamic range image cubes and can work with an ARC node to apply self-calibration to the data. For the use of nonstandard analysis

techniques and the utilisation of the advanced algorithms and software developed by a specific ARC node, ALMA users may also arrange a visit to this ARC node, where the best possible support can be provided.

The ARC nodes provide the latest (or any relevant) version of CASA, installed on high performance workstations with ample disk space. For any first line CASA support, please visit the CASA Guides website at <http://casaguides.nrao.edu>.

Additional help from the European ARC

Archive research support

A complete, synchronised mirror of the ALMA archive is kept at ESO and is a valuable resource for data mining. Assistance for archive research encourages broader approaches to scientific investigations. Therefore, scientists in the ARCs and ARC nodes provide support for astronomers accessing the ALMA science archive. The ALMA Science Archive is available through the User Portal.

Prospective ALMA users are likely to want to consult the Archive when planning observations as well as for data mining. To ensure that the archive (and ALMA) is exploited to its full potential, ARC node or CoE staff can assist

users in using the archive. This includes face-to-face visits. Note that only PIs and Delegates can access projects completed within the last 12 months.

The ARC nodes and CoE can also help users extract information effectively (such as potential calibration source properties, or whether there are public domain data matching the required criteria) as well as assist with analysis of archive data.

*Basic archive research support: use the [Helpdesk](#).
Advanced archive exploitation support: visit an ARC node (use the [Helpdesk](#) to coordinate a visit).*

Need training?

The European ARC regularly organises workshops, tutorials and schools related to ALMA. Dedicated CASA data reduction tutorials are held throughout Europe, as well as sessions describing the ALMA Observing Tool and more general workshops explaining the current ALMA capabilities. The EU ALMA newsletter and announcements provide a list of all scheduled activities and the mailing lists of the ESO ARC and the local nodes are used to disseminate this information.

Need information on upcoming ALMA related tutorials and workshops? Check the [EU ARC webpages](#) or the [EU ARC announcements](#), or subscribe to the mailing list of your local ARC node.

General questions

For any general questions related to ALMA capabilities, availability of observing modes, or other relevant topics, please use the Helpdesk.

Any questions related to ALMA: use the [Helpdesk](#)

How to organise a face-to-face visit

The preferred way to obtaining any type of support is through the ALMA Helpdesk (select “Submit a Ticket”, then “Face to Face Support (EU)”). In the Helpdesk ticket and in order to make the most of the visit to an ARC node or CoE, the user should give as much information as possible on the purpose of the visit and the proposed node and dates.

Together with the staff at the ARC network, it is decided which node the user should visit and on which dates. Support staff at this node are responsible for arranging the details of the visit. Each visitor is assigned a dedicated member of staff for support purposes and can expect that this support person will be able to respond as quickly as possible to their requests for help, within reason.

Users normally visit their local node. If there is no ARC node associated with the country where the user is working, or if the user and the ARC (node) staff decide that specialist support is best provided by another node, a visit to another node can be organised.

ALMA data reduction will normally be performed using dedicated computing facilities such as those provided by the ARC nodes. The data volumes are too large for laptops in most cases, except for the later stages of image or spectral analysis, possibly using subsets of the data. Please consult ARC staff if you want to use your laptop.

To organise a face-to-face visit: use the [Helpdesk](#)

To ensure data are downloaded locally and ready for the PI, the PI must give explicit permission for processed and raw data to be downloaded from the ALMA Science Archive by the ARC node staff via the “data delegation” tab in their ALMA

Science Portal User Profile. ARC nodes provide facilities for the user to copy data onto a hard disk drive. Data backup should be discussed prior to the visit, so that the user can bring appropriate storage devices. Some nodes can also support ftp of data provided the link at the user's home institute is fast enough.

ARC nodes keep a copy of the data the user has been working on for at least a few weeks after the end of the user's visit. This enables the node staff to address any post-visit questions which may arise.

After a visit, the user is requested to submit a feedback form on the service received, indicating, among other things, whether the goals of the visit were met.

Funding for visits

It is foreseen that most users will apply for funding from their local agencies for travel to one of the ARC nodes or CoE. Some ARC nodes have access to funds for visiting scientists. Please check the details in the next section, or visit the ARC node web pages. Funding is also available through the RadioNet MARCUs (Mobility for ARC Users) networking activity. Requests for MARCUs funding should be sent via [this form](#). MARCUs-funded visits are paid by the user's institute, which then is reimbursed by RadioNet. Note that filling the feedback form is mandatory for users supported by MARCUs funding.

If you receive funding from RadioNet MARCUs and/or useful assistance from a European ARC node or the ESO ARC, please include this in the Acknowledgement section of any paper based on ALMA data of the supported project.

The ARC nodes and CoE

Each node has its own area(s) of expertise. Users are encouraged to visit the individual ARC node web pages to obtain the most up-to-date information. Most ARC nodes also maintain email lists that are used to inform their local communities about activities and region-specific news items.

The Italian node in Bologna

The Italian ARC is hosted by the Osservatorio di Radioastronomia in Bologna. The node currently employs three staff members and four postdocs, experts in mm-astronomy and/or interferometry, and has the support of a dedicated system manager. Support is offered with the use of ALMA tools for proposal preparation and submission, tracking of accepted ALMA projects, data reduction with CASA, ALMA archive mining, and the special area of polarimetry. This is mostly done via face-to-face visits or Helpdesk tickets.

The Italian ARC is also developing new CASA tasks, and investigates new techniques of data reduction and handling and collaborates on the ALMA archive development. It organises seminars, tutorials and workshops to inform and train the Italian community and to stimulate scientific discussion and collaborations. ARC personnel is involved in teaching and supervising students.

The Italian ARC researchers are involved in several (sub-)mm scientific projects covering different topics (from galactic and local-Universe studies to high-redshift galaxies and cosmology), and offer general scientific support on their specific area of expertise for continuum and spectroscopic observations.

Visit the Italian ARC node at www.alma.inaf.it

At the host institute, the ARC node offers fully equipped office space where it can simultaneously accommodate two (or more) visitors for face-to-face support and visits. Visitors are provided with desktop computers (or laptop sockets), directly connected to a dedicated computer server, to access the data and run the CASA data reduction software.

The ARC node is connected to the outside world through a high-speed optical fibre network, allowing fast data transfer (10 Gbit/sec). ALMA and CASA users can access the ARC node cluster facilities and 1 TB of disk space for a period of six months after their face-to-face visit. On request, access can be extended to longer periods.

No local funding is available for visitors but eligible visitors can apply for MARCUs support.

The Dutch node in Leiden: Allegro

Allegro, the ARC node in the Netherlands, is located at Leiden Observatory. The node employs five to six postdocs and two staff members who coordinate the node's activities. Allegro provides general face-to-face support. Allegro also offers expert advice on high-frequency observing (roughly speaking bands 8, 9, and 10) including proposal planning, setup of observations and calibration strategies, reduction techniques, and data analysis. In

Visit the Dutch ARC node at www.alma-allegro.nl

addition, Allegro offers help in using radio-techniques for wide-field imaging and mmVLBI, and hosts a large suite of data analysis tools that can be consulted on-line or in person. Allegro has guest desks available for visitors, who can make use of the dedicated high-performance computing facilities and connected fast-access storage system.

The German node in Bonn/Cologne

The German ARC node is a collaboration between the astronomical institutes at the universities of Bonn and Cologne. The node's main facilities are located at the Argelander-Institut für Astronomie in Bonn, right next to the Max-Planck-Institut für Radioastronomie and a 30-minute commute from the University of Cologne. Four staff and seven postdocs are associated with the ARC node, working on various support tasks. Areas of special expertise include the combination of interferometric data with single-dish data, polarimetry, and advanced analysis and modeling of spectral line data. The node maintains and continuously extends the Cologne Database for Molecular Spectroscopy and offers a number of data analysis tools developed in-house. In close collaboration with the Max-Planck-Institut für Radioastronomie, the node also contributes to the efforts towards integrating ALMA into global mm-VLBI networks.

Visitors to the German ARC node can use the spacious ARC visitors' room equipped with workstations, switches for laptops and a conference phone, and a parent-child office located next to the ARC visitors' room. The workstations have fast connections to two high-performance servers that are reserved for ALMA-related work and run all the latest ALMA software necessary for the proposal stage, data reduction and analysis and for ALMA archive searches. Up to two visitors can be accommodated simultaneously.

Visit the German ARC node at www.astro.uni-bonn.de/ARC

In addition to user support, development and services to ALMA operations, the German ARC node reaches out to the local community with annual community meetings, a local newsletter and talks at seminars and conferences. The node's periodic training opportunities include (remote) participation in a comprehensive hands-on lecture series on radio interferometry, a CASA tutorial and ALMA-OT tutorials. In addition, some members of the ARC node are active in general public outreach.

No local funding is available for visitors but eligible visitors can apply for MARCUs support.

The IRAM node in Grenoble

IRAM (Institut de Radio-Astronomie Millimétrique) is an international institute supported by the Max Planck Society (Germany), the CNRS (France), and the IGN (Spain). The institute is operating the 30-m Pico-Veleta antenna and the Plateau de Bure mm-interferometer. The construction of NOEMA, the successor to the Plateau de Bure observatory, has now started. The IRAM ARC node activities are built on the in-house experience with user support and on the involvement of the institute in the ALMA design and construction phases (in particular the development of the ALMA real-time Telescope Calibration software).

The face-to-face support for ALMA will use the same procedures and infrastructures as those used for the Plateau de Bure support. A computer room dedicated to data reduction is available and each project is assigned a local contact. A total of six astronomers are associated with the ARC node.

The IRAM ARC node provides a good opportunity to use synergies between the IRAM instruments and ALMA, as, e.g., an access to the full sky or to large-scale surveys with the 30-m. In addition to general face-to-face support, specific domains of expertise at IRAM include the calibration of (sub-)mm interferometric data or wide-field imaging and the inclusion of short-spacings.

Visit the IRAM ARC node at www.iram.fr/IRAMFR/ARC

Funding is available for travel and accommodation, with the usual IRAM rules: one mission/one person per project. Funding is however limited to astronomers affiliated to one of the IRAM funding agencies.

The Nordic node in Onsala

The Nordic ARC node's main mission is to support ALMA users in Sweden and other Nordic and Baltic countries. The node is physically located at Onsala Space Observatory (OSO) near Gothenburg, Sweden. OSO is the Swedish National facility for Radio Astronomy and is hosted by Chalmers University of Technology. OSO has a long history in mm/submm astronomy and presently runs the 20-m single dish telescope at Onsala and is a

partner in the APEX submillimetre single dish telescope at the ALMA site in Chile.

Visit the Nordic ARC node at www.nordic-alma.se

The ARC node has high capacity data links for downloading archival data and high performance computer facilities to support ALMA data reduction. There is ample dormitory and guest room accommodation on-site to support data reduction visits and small workshops. The ARC node currently employs four permanent staff members, five postdocs and IT support. The ARC node includes experienced millimetre wave observers, and specialists in interferometry algorithm development.

The node is active in community outreach to promote ALMA in the Nordic countries. The node also organises meetings for the whole Nordic community to explain the science that can be done with ALMA and train users to plan, submit, and reduce ALMA projects. As well as giving general face-to-face support, the node has expertise in the areas of astrometry, robust phase self-calibration, multi-frequency synthesis, deconvolution algorithms and polarisation. The node is also involved in the study for phasing up ALMA for VLBI and develops and maintains several advanced data analysis packages. In collaboration with academic staff at Chalmers, there is also extensive scientific expertise on-site in the areas of radiative transfer and astrochemistry.

No local funding for travel and accommodation is available but eligible visitors can apply for MARCUs support.

The UK node in Manchester

The UK ARC Node is centrally located at the University of Manchester, Jodrell Bank Centre for Astrophysics, conveniently placed for all national and international transport links and high-speed internet connectivity. Five ARC staff (plus IT and administrative support) divide their time between face-to-face user support and related research and development. Facilities for visiting astronomers include a large, dedicated room with ample desk space and powerful computers, which can also be used for small meetings and tutorials. Absolute beginners or experts are welcome and resources can be made available by arrangement for teams working on ALMA projects with demanding computational and skills requirements.

Visit the UK ARC node at <http://www.alma.ac.uk>

The UK ARC Node draws on local and national expertise to offer specialised assistance in a number of areas:

- Interferometry skills, in particular advanced calibration, extended array configurations (high resolution), wide-band high-sensitivity continuum imaging (multi-frequency synthesis), mosaicing and wide-field imaging (including combining ALMA configurations to improve image fidelity), and polarimetry.
- Multi-instrument research, in particular comparing data from different telescopes, (e.g. Herschel, JCMT, e-MERLIN, EVN) combining ALMA and compatible data, and interoperability between CASA and AIPS. The UK ARC Node is part of the JBCA Interferometry Centre of Excellence which provides support for these instruments, others such as LOFAR and prospective SKA users.
- Data management, in particular data mining the ALMA Archive, development of pipelines and scripts for large projects, standards and documentation for data publication.
- Simulations using the OST (<http://almaost.jb.man.ac.uk/>) or inside CASA: choosing suitable input images, estimating the brightness on ALMA scales and so on.

STFC PATT funds can now be used to fund visits from UK institutes to the UK ARC Node. If your institute does not hold a PATT linked grant, you can apply for support through the usual PATT process, see <http://www.stfc.ac.uk/1506.aspx>

The Czech node in Ondrejov

The Czech node is located at Ondrejov Observatory near Prague. The node has available an office for three visitors equipped with computers dedicated to ALMA. Affordable accommodation in the guest rooms of the Astronomical Institute at Ondrejov can be offered. The node's duties are currently distributed over two staff members, two postdocs, and one technician.

The team is ready to provide support in preparation and submission of proposals to ALMA, with interpretation of observations, with some numerical modelling of

Visit the Czech ARC node at www.asu.cas.cz/alma

the observed processes, and interstellar molecular line identification. The expertise areas of the Czech node are:

- Solar ALMA observing mode: specifics of proposal preparation, calibration and imaging; solar radio spectroscopy; non-LTE radiative transfer; physics of the solar chromosphere, prominences and filaments; solar flare plasma astrophysics.
- Galaxy evolution in clusters and groups, triggered star formation, proto-planetary disks, Galactic Centre and AGN.
- High resolution microwave laboratory spectroscopy of small molecules with a focus on hyperfine effects, spectroscopy of unstable transient molecular species, theoretical analyses of molecular energies, and transition (dipole) moment measurements.

No local funding is available for visitors but eligible visitors can apply for MARCUs support.

The Portuguese CoE in Lisbon: PACE

The Portuguese ALMA Centre of Expertise (PACE), operating since July 2014, is hosted by the Institute of Astrophysics and Space Sciences and located at the Astronomical Observatory of Lisbon. Two staff members and six postdocs devote part of their time to PACE duties, such as organisation of events for community support, or face-to-face user support on proposal preparation and submission, as well as archival research.

Visit the Portuguese CoE at
<http://pace.oal.ul.pt>

Dedicated workstations, including access to a Nonius Cluster, are available to visitors.

The PACE astronomers have experience in mm-to-radio science. Their interests lie in the fields of galaxy formation and evolution, planetary atmospheres and archival research. The experience in these topics will be strengthened in order to improve the capabilities of the EU ARC network user support.

PACE also maintains a strong outreach team through which it educates the national community on mm-to-radio astronomy.

No local funding is available for visitors but eligible visitors can apply for MARCUs support.



The Atacama Large Millimeter/submillimeter Array (ALMA), an international astronomy facility, is a partnership of the European Organization for Astronomical Research in the Southern Hemisphere (ESO), the U.S. National Science Foundation (NSF) and the National Institutes of Natural Sciences (NINS) of Japan in cooperation with the Republic of Chile. ALMA is funded by ESO on behalf of its Member States, by NSF in cooperation with the National Research Council of Canada (NRC) and the National Science Council of Taiwan (NSC) and by NINS in cooperation with the Academia Sinica (AS) in Taiwan and the Korea Astronomy and Space Science Institute (KASI).

ALMA construction and operations are led by ESO on behalf of its Member States; by the National Radio Astronomy Observatory (NRAO), managed by Associated Universities, Inc. (AUI), on behalf of North America; and by the National Astronomical Observatory of Japan (NAOJ) on behalf of East Asia. The Joint ALMA Observatory (JAO) provides the unified leadership and management of the construction, commissioning and operation of ALMA.

